Jupiter flash

August 7th 2019, flash at 04:07:30 UTC



Color image at 04:10:36 UTC

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Observations with a Celestron 8" telescope from Cibolo, Texas

A new small impact in Jupiter in August 7th, 2019

Ricardo Hueso (UPV/EHU), Marc Delcroix (SAF) Ethan Chappel (discoverer of the impact) Agustín Sánchez-Lavega (UPV/EHU)

Ramanakumar Sankar (FIT) Csaba Palotai (FIT)

Jon Juaristi-Campillo (UPV/EHU & Heidelberg University)

& The Impact Recovery Team







Small impacts in Jupiter - Common phenomena but frequency unknown

6th Impact "flash" found in Jupiter since 2010 (2 flashes in 2010, 1 in 2012, 1 in 2016, 1 in 2017) Hueso et al. ApJ, 2010, Hueso et al. A&A, 2013

Hueso et al. A&A, 2018: "Small impacts in the Giant planet Jupiter" A report of all previous flash impacts in Jupiter. Impact rate predicted: low (10–65 impacts per year) with only 4–25 observable impacts per year



First impact found by software and not the observer. The impact was found using DeTeCt (software by M. Delcroix, J. Juaristi & R. Hueso)

The 2019 impact: Very clear signal ! Bigger than the impact in 2010 and smaller than the brightest flash in 2012

Impact discovery image



Image generated with DeTeCt

Absence of debris on amateur images

Impact Time=2019-08-07T04:07:30 UT



Green image at Impact time +3min.



Methane image at Impact time +28min.



Many other images acquired by Ethan Chappel between these two images. None of them shows any possible debris.



DeTeCt: A software tool in windows to automatically analyze videos of Jupiter

EPSC-DPS2019-970 : Jupiter and Saturn impact detection project, Delcroix et al., Monday 14:25



3 impacts in 127 days of accumulated data: 8.6 "detectable" impacts per year in the visible side of Jupiter or ~ 20 impacts per year counting impacts in the rest of the planet

This number does not consider problems related with seeing or video quality and the real number should be larger We can increase our last estimation of impact rate in Jupiter from **10-65 to 20-65 impacts per year**.

DeTeCt project goals: (1) That observers run DeTeCt after their observing sessions regularly

(2) That prominent observers with dozens of Terabytes of data stored in their hard-drives use the software to find "past" impacts. Kind support of Clyde Foster (South Africa) and others

The size and energy of the 2019 impact

Automatic pipeline analysis using correlation of images and aperture photometry masks









Not the first case with fragmentation but the first one we try to model!



Less clear fragmentation in the light-curve of the 2017 impact

Hueso et al. *A&A* (2018)

The size and energy of the 7th August 2019 impact



Astro-Physics Advanced Convertible Barlow (considered negligible)

Calculation of kinetic energy by Ramanakumar Sankar (Florida Institute of Technology)

From light-curve: 8.9×10^{12} J in red From black-body temperature~6500 K \rightarrow ~ 8.6×10^{13} J total luminous. Luminous efficiency ~ 8.5%Total energy ~ 1.0×10^{15} J (240 kT) 50% of the Chelyabinsk meteor impact in 2013 Almost exactly the same energy as estimated from a scale comparison with the 2012 impact by R. Hueso

Estimated mass ~4.4 - 5.5x10⁶ kg Maximum diameter: 16 m (if cometary density p=0.25 kg/m3)

Not large enough to leave a traceable debris

(largest SL9 fragment with a debris SL9-N had an estimated size of 47 m, 25 times more massive)

This object is slightly smaller than the biggest flash observed in Jupiter (the September 2012 impact) which partially saturated the detector and had a maximum mass of 1000 Tn and a maximum size estimate of 20 m in diameter.

Models of impact fragmentation (also courtesy of Ramanakumar Sankar and Csaba Palotai at FIT)



Simulations based on Borovicka et al. (2007) fragmentation model.

The code starts with initial conditions of the object (mass, velocity, angle, height) and a prescription of the points of fragmentation (time and duration in the light-curve, total mass loss in each fragment and number of fragments. Inputs include the type of material (ablation coefficient, shape density coefficient).

Simulation above: 0.45 kTon object impacting at 69 km/s with an angle of 25^o.

This fit of the light-curve requires **high density** (4.5 g/cm3; **stony-iron meteor**). Terminal depth ~ 15 km below the point the bolide ignites (possibly at z=110-120 km) **Terminal depth** ~ z=95-105 km or ~6-8 mbar Diameter of this stony-iron object: 12.5 m

<u>Non-unique solution</u>. Currently the team at FIT is exploring the space of parameters to rule out a possible cometary impact.

Another impact alert in 2019: August 30 (Netherlands) Visual observation

Visual observation of a flash in Jupiter near sunset with Jupiter at 15^o elevation. Telescope: C14, most convincing report so far (except the one by William Petersen in 2012 that led to the discovery of the September 2012 impact by George Hall both in USA) of the characteristics of an impact flash,

Only information available: August 30 18:15-18:30 UT) & Location over Jupiter (SEB)



An Impact Recovery Team

Observers close in time:	Michel Jacquesson, (France)	18:52-19:37	'UT
	Armando Vaccaro (Italy)	18:59-19:22 UT	
	Javier Beltrán Jovani (Spain)	19:36	UT

Observational alert in Italy, Malta, Greece, Romania & South Africa (Jupiter at their peak elevation around 18-19 UT) Help from M. Vedovato, Manos Kardasis, Ioannis Bouhras, Constantin Sprianu, Clyde Foster and Alexei Pace



No success in finding a video observation of the planet close to the right time. Larger alert released on Sept. 9 to observers in African countries with help of Salma Sylla and Clyde Foster. No impact scar on Jupiter in images obtained the same and later nights.

Estimated impact rate in Jupiter

Hueso et al. A&A 2018



The new impact increases slightly the lower limit of the estimated impact rate of superbolides but does not change things significantly.

WHAT DO WE NEED?

1) Better estimate of objects masses.

Requires better knowledge of flash brightness temperature, a new flash observed simultaneously with two filters

2) Better understanding of the more frequent impact rate of smaller objects Requires dedicated observing programs with professional telescopes

3) Better understanding of the less frequent impact rate of larger objects Examen of accumulated observations on telescopes such as HST.

Survey efficiency for HST and amateur searches of impacts



Efficiency to discover a SL-9 N-like debris field (47-m object) by random chance over HST images: ~3%.

Need at least an impact rate of 7 objects per year of this size to produce a high probability of discovery.

Estimates in Hueso et al. (2018) are 1-3 objects per year of this size. It would be very difficult to find impact debris on HST images by chance.

Efficiency to discover a 2009 impact debris field (~150 m object) by random chance over HST images: ~13%.

Need at least an impact rate of 1-2 objects per year of this size to produce a high probability of discovery.

Estimates in Hueso et al. (2018) are 0.2-0.4 objects per year of this size.

Observational surveys for small and big impacts at:

Hampton University (Kunio Sayanagi & Benito Loyola) ~300 hrs per year of observations with 8-14" telescopes

Senegal (Salma Sylla) Excellent near equatorial location with the capability to provide long observing sessions

Pic du Midi, France (François Colas) ~ A few tens of hours 15-50? per year with a 1m telescope in a site with excellent seeing and optics

Calar Alto Observatory, Spain (R. Hueso & A. Sánchez-Lavega) ~ 15 hrs per year at a 2.2 m telescope with two simultaneous cameras

Several surveys by amateurs: Most dedicated amateurs (A.Wesley, C. Foster, ...) can observe Jupiter ~300 hrs per year

Jupiter System Dynamics Observatory at Sun-Jupiter Lagrangian Point One (PI:H.W. Hsu, University of Colorado)

Dedicated instrument in the mission proposal. Mission concept proposal under study at NASA

Targeting impact scars caused by bigger impacts: One every 4-10 years an object could leave a traceable debris in Jupiter's atmosphere

✓ HST Cycle 27 ToO proposal accepted (PI: Imke de Pater)

✓ Possibility to detect stratospheric exogenous species with ALMA from "small" impacts

✓ JWST Cycle 1 opportunities